#### A decaying dark matter model for 3.5 keV X-ray line

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based on SB, 1410.1992

#### Outline

- · Zee-Babu neutrino model
- Extended Zee-Babu model including DM
- 3.5keV X-ray line signal in extended Zee-Babu model
- Conclusions

# Zee-Babu model for neutrino masses

• Zee-Babu model for Majorana neutrinos: two new charged scalars  $h^+$ ,  $k^{++}$  with L=-2 are introduced

 $\mathcal{L}_{Y} = f_{ab}(\psi_{aL}^{Ti}C\psi_{bL}^{j})\epsilon_{ij}h^{+} + h'_{ab}(l_{aR}^{T}Cl_{bR})k^{++} + \text{H.c.}$ 

 $\mathcal{L}_{h-k} = -\mu h^+ h^+ k^{--} + \text{H.c.}$  L-violating soft term

 $(\mathcal{M}_{\nu})_{ab} = 8 \mu f_{ac} m_c h^*_{cd} m_d f_{db} I_{cd}$ 



- Small neutrino masses are due to loop suppression
- Testable at LHC contrary to seesaw mechanism
- We extend the model to local B-L symmetry and introduce B-L charged DM candidate

#### The model

$$i = 1, 2, 3$$

Fields	$q_i$	$\ell_i$	$h^+, k^{++}$	$\varphi$	$\eta$	$N_{R_i}$	$\psi_{m{i}}$
B-L	1/3	-1	2	2	0	-1	1/3, 7/3, 13/3
$Z_2$	+	+	+	+	+	_	±

- $\psi$  (Dirac DM) generate transition MD op.
- $\varphi$  U(1)<sub>B-L</sub> breaking scalar
- $\eta$  Light scalar for relic density & small scale problems
- $N_R$  (Majorana) cancel gauge anomaly

#### The model



- $LHN_R$  forbidden by Z<sub>2</sub>
- $LH\psi$  forbidden by B-L charge assignment



: dominant contribution for neutrino masses

#### Local U(1)<sub>B-L</sub> symmetry + $(\lambda_{\mu}\varphi k^{++}h^{-}h^{-}+h.c) \longrightarrow \mu k^{++}h^{-}h^{-}$

dynamically generates LV µ-term

$$\mathcal{L}_{\Psi} = \overline{\psi_i} i \gamma^{\mu} D_{\mu} \psi_i - m_{\psi_i} \overline{\psi_i} \psi_i - f_{12} \Big( \overline{\psi_1} \psi_2 arphi^* + \overline{\psi_2} \psi_1 arphi \Big) - f_{23} \Big( \overline{\psi_2} \psi_3 arphi^* + \overline{\psi_3} \psi_2 arphi \Big)$$

ψ's have "flavor" changing Z' and φ interaction <φ>: U(1)<sub>B-L</sub> → Z<sub>6</sub> :not broken by quantum gravity-> guarantees absolute stability of DM

$$\mathcal{L}_{N_R} = \overline{N_{R_i}} i \gamma^{\mu} D_{\mu} N_{R_i} - \frac{1}{2} \Big( \lambda_{N_{ij}} \varphi \overline{N_{R_i}^c} N_{R_j} + h.c. \Big)$$

N<sub>R</sub>'s not DM candidates. Can decay through M<sub>Pl</sub> suppressed Op.

$$\lambda_{ij}\ell_i H N_{R_j}, \quad \frac{1}{M_{\rm Pl}}\ell_i H N_{R_j}\eta, \quad \frac{1}{M_{\rm Pl}^2}N_{R_i}\ell_j\ell_k\bar{e}_l, \quad \frac{1}{M_{\rm Pl}^2}N_{R_i}\bar{d}_j\bar{d}_k\bar{u}_l,$$

#### Decaying DM for 3.5 keV X-ray

 If the decay can be described by transition dipole moment operator (TMDO)



Through two-loop Barr-Zee

$$\mathcal{L}_{\rm eff} = \frac{1}{\Lambda} \overline{\psi_1'} \sigma_{\mu\nu} \psi_2' F^{\mu\nu}$$

is generated

$$\gamma \\ h^+, k^{++}$$
  
 $Z' \qquad \phi \\ \psi'_2 \qquad \psi'_i \qquad \psi'_1$ 

$$\begin{split} \frac{1}{\Lambda} &\simeq \sum_{s=h^+,k^{++}} \frac{8eg_{Z'}^2 \Delta Q_{\psi} Q_s Q'_s \lambda_{\varphi s} \delta^2 \cos 2\theta_{12} s_{13} s_{23}}{(4\pi)^4} \\ &\times \int_0^1 dx \int [d\beta] \frac{x \beta_4^2 m_{\psi_3'}^2}{\left(\beta_1 m_{Z'}^2 + \beta_2 m_{\phi}^2 + \beta_3 m_s^2 / (x(1-x)) + \beta_4^2 m_{\psi_1'}^2\right)^2}, \end{split} \delta = \Delta m_{31} / m_{\psi_3'}$$

#### Kinetic mixing

- Z'- $\gamma$  mixing:  $-\sin\chi \hat{Z}'_{\mu\nu}\hat{F}^{\mu\nu}/2$
- Tree-level TMDO? No.
- Canonical kinetic enegy from non-unitary
  - $\begin{pmatrix} \hat{A} \\ \hat{Z}' \end{pmatrix} = \begin{pmatrix} 1 \tan \chi \\ 0 & \sec \chi \end{pmatrix} \begin{pmatrix} A \\ Z' \end{pmatrix} \qquad \qquad \Delta \mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} \frac{1}{4} Z'_{\mu\nu} Z'^{\mu\nu} + \frac{1}{2} M_{Z'}^2 Z'_{\mu} Z'^{\mu},$
- The transformation is unique for massive Z'
- Covariant derivative

$$egin{aligned} D_\mu &= \partial_\mu + i \hat{e} Q \hat{A}_\mu + i \hat{g}_{Z'} Q_{Z'} \hat{Z}'_\mu \ &= \partial_\mu + i \hat{e} Q A_\mu + rac{i}{\cos\chi} (\hat{g}_{Z'} Q_{Z'} - \hat{e} Q \sin\chi) Z'_\mu. \end{aligned}$$

 $\cdot\,$  Photon couples only to EM-charge

#### Fine-Tuning

- Small parameters in the model:  $\Delta m_{12} << m_{\psi}$
- 't Hooft naturalness criterion "A parameter is naturally small if setting it to zero increases the symmetry of the theory"
- Setting  $\Delta m_{12}$  to zero increases the symmetry U(2)flavor symmetry in ( $\psi_1, \psi_2$ ) flavor space
- Small  $\Delta m_{12}$  is technically natural

To explain 3.5 keV X-ray line, we need

$$\Lambda = (6.94 \times 10^{14} - 2.95 \times 10^{15}) \left(\frac{m_{\psi'_2}}{\text{GeV}}\right)^{-1/2} \text{GeV}.$$

- In our case it is obtained not by heavy particle but by loop suppression
- Benchmark point

$$\delta = 0.2, \ \theta_{12} = \theta_{23} = 0.2, \ m_{\phi} = m_{h^+} = m_{k^{++}} = 1 \text{ TeV}$$
  
 $\lambda_{\varphi h} = \lambda_{\varphi k} = 1. \qquad m_{\eta} = 1 \text{ MeV}$ 

• Constraints: LUX DM direct search, perturbativity, DM relic abundance, LFV,  $\frac{M_{Z'}}{g_{Z'}} > 7 \,\text{TeV}$ 



#### The role of $\eta$

- $\cdot \,\,\eta$  gives the correct relic abundance
- Light  $m_{\eta}$  (1—10 MeV) can enhance  $\psi'_{1(2)}, \psi'_{1(2)} \rightarrow \psi'_{1(2)}, \psi'_{1(2)}$ and can explain small scale structure problems, corevs-cusp and too-big-to-fail problems, if

$$\sigma_T / m_{\psi_1'} \sim 0.1 - 10 \,\mathrm{cm}^2 /\mathrm{g},$$



• For  $y_i \sim \mathcal{O}(1)$ , the self-interaction occurs in the non-perturbative  $(\alpha_y m_{\psi'}/m_\eta \gtrsim 1 \text{ with } \alpha_y \equiv y^2/4\pi)$ and classical  $(m_{\psi'} v_{rel}/m_\eta \gg 1)$  regime



### Conclusions

- 3.5keV X-ray line signal can be explained in extended Zee-Babu model with DM
- Some parameter region is sensitive to the next generation DM direct search
- Light η achieves the correct relic abundance and also solves the small scale structure problems, core-vscusp and too-big-to-fail problem
- h<sup>+</sup>, k<sup>++</sup> may be discovered at LHC